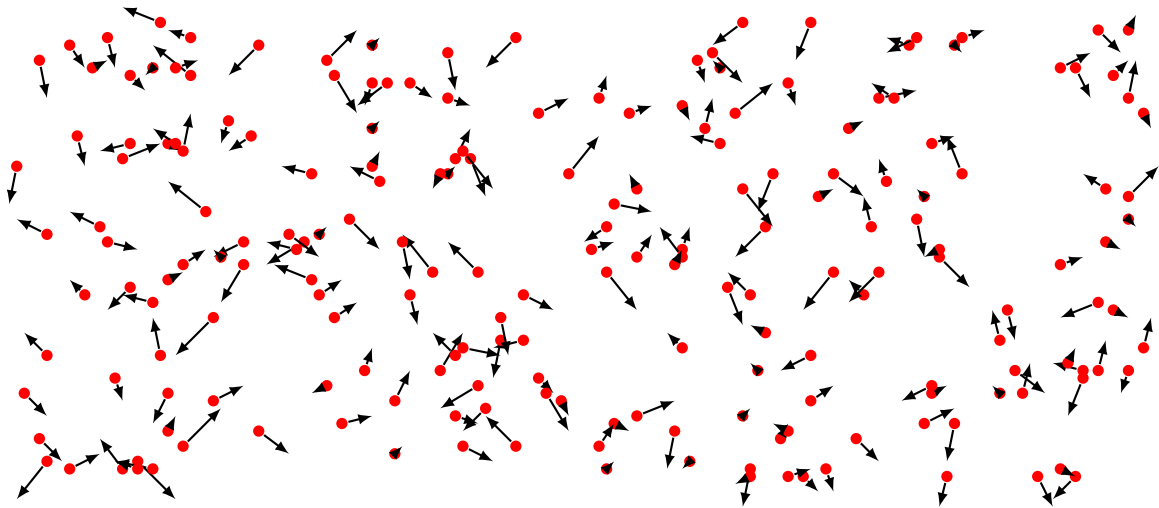


Section 2

Signal and noise

Randomness



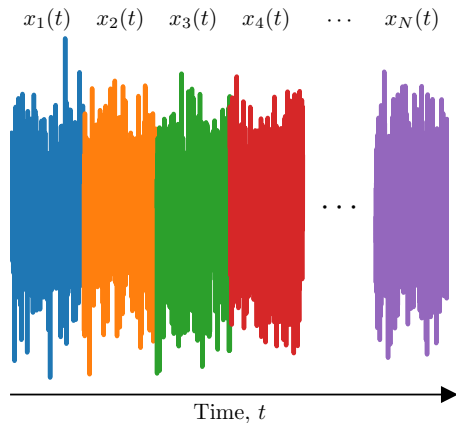
Random noise and the ensemble average

Consider a (long) noise time series $x(t)$.

We can divide it into N segments:

$$x_1(t), x_2(t), \dots, x_N(t).$$

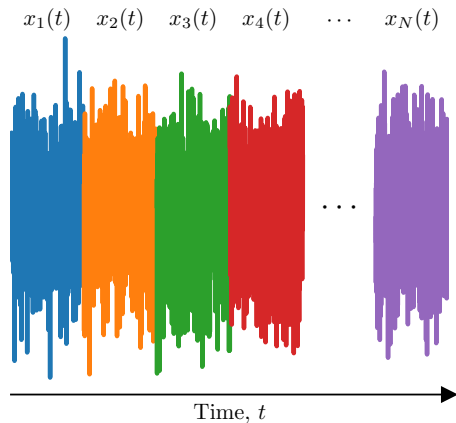
The collection of these segments is an **ensemble** of noise realizations.



The ensemble average

Noise is random, but its **ensemble-averaged properties** are deterministic.

For example, the mean value of each realization $\bar{x}_1, \bar{x}_2, \dots, \bar{x}_N$ is random. However, its ensemble average $\langle \bar{x} \rangle$ has a single, well-defined value.



Given an ensemble of time series $x_1(t), x_2(t), \dots, x_N(t)$, each of length T , the *power spectral density* is the ensemble average of the normalized square magnitude of the Fourier transforms:

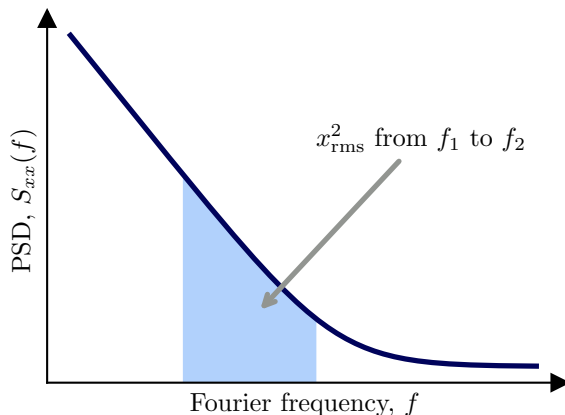
$$S_{xx}(f) = \frac{2}{T} \langle |\tilde{x}(f)|^2 \rangle,$$

where T is understood to be arbitrarily large. (The factor of 2 will be explained later.)

The quantity $\sqrt{S_{xx}(\Omega)}$ is called the *amplitude spectral density*.

Why power spectral density?

Most random noises have fluctuations at all frequency scales. Power spectral density tells us how to compute the expected mean-square noise power occurring between any two frequencies:



In classical mechanics, $x(t)$ is real. Consequently, $\tilde{x}(f)$ satisfies $\tilde{x}(-f) = \tilde{x}^*(+f)$.

Therefore, negative frequencies are redundant. However, they still contain half of the total noise power.

Experimentalists usually work with a single-sided spectral density $S(f)$, defined for $f \in [0, \infty)$, and a factor of 2 is included to account for the power in negative frequencies.

Theorists sometimes work with the double-sided spectral density, defined for $f \in (-\infty, \infty)$.

Unfortunately, papers are not always explicit about which convention they use!

Two spectral densities to commit to memory

The voltage fluctuations across ideal resistor with resistance R at temperature T has single-sided power spectral density

$$S_{VV}(f) = 4k_{\text{B}}TR \quad (k_{\text{B}} = \text{Boltzmann's constant}).$$

The optical power fluctuations due to photon shot noise in a beam of light of frequency ν and power P has single-sided power spectral density

$$S_{PP}(f) = 2h\nu P \quad (h = \text{Planck's constant}).$$

Given two random noises x and y , their **cross-spectral density** is

$$S_{xy}(f) = \frac{2}{T} \langle \tilde{x}^*(f) \tilde{y}(f) \rangle.$$

The cross-spectral density in general is complex-valued.

For random noises x and y , their **magnitude-squared coherence** (often just **coherence**) is

$$|\gamma|^2 = \frac{|S_{xy}(f)|^2}{S_{xx}(f)S_{yy}(f)}.$$

By the Cauchy–Schwartz inequality, $0 \leq |\gamma|^2 \leq 1$.

When $|\gamma|^2 = 0$, the two noises x and y are completely uncorrelated. When $|\gamma|^2 = 1$, the two noises are completely correlated.

Pay attention to how the ensemble averaging is done:

$$|\gamma|^2 = \frac{|\langle \tilde{x}^*(f)\tilde{y}(f) \rangle|^2}{\langle \tilde{x}^*(f)\tilde{x}(f) \rangle \langle \tilde{y}^*(f)\tilde{y}(f) \rangle}.$$

Data records these days are discrete. So is the Fourier transform:

$$\tilde{x}(f_j) = \sum_{i=0}^{N-1} x(t_k) \exp(-2\pi i f_j k \Delta t).$$

(This follows the `numpy` conventions.)

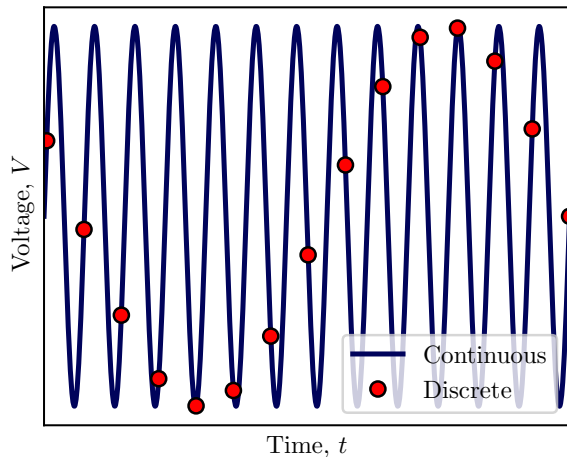
Note that $N = T \Delta t$, and that

- Frequency points are separated by $\Delta f = 1/T$ (the sampling frequency).
- The highest frequency is $f_{\text{Nyq}} \leq 1/(2\Delta t)$ (the *Nyquist frequency*).

Complications of discretisation: aliasing

Signal content above the Nyquist frequency is aliased below the Nyquist frequency.

The analog signal must be filtered (anti-aliased) before being digitised to remove this frequency content.

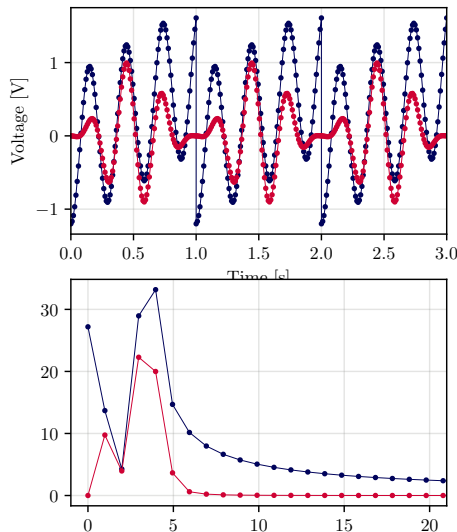


Complications of discretisation: periodicity

The DFT operates on a periodic domain.

Discontinuities at the ends of a finite time series appear as (usually unwanted) features in the DFT. (Generically, step-like functions in the time domain produce Laplace-domain functions like $1/s$.)

Most spectral density estimation algorithms **window** or **taper** the time series to zero at either end to reduce this effect. See `scipy.signal.windows` for some common ones.



Implementation

The “standard” method of Welch’s overlapping periodograms:

```
freq_arr, csd_arr = scipy.signal.csd(  
    x, # Time series x(t)  
    y, # Time series y(t)  
    fs=<val>, # Sampling frequency of time series (in hertz)  
    window="hann", # Hann is a nice general-purpose window function  
    nperseg=<val>, # Number of points in each segment  
    noverlap=<val>, # Defaults to nperseg // 2, good choice for Hann window  
    nfft=None, # Zero-pads data. Never use this: leave it as None.  
    detrend="linear", # Usually better than "constant"  
    return_onesided=True, # The single-sided (experimentalist) convention  
    scaling="density", # Scales to spectral density units  
    average="mean",  
)
```

Section 3

Feedback control

Feedback on a harmonic oscillator

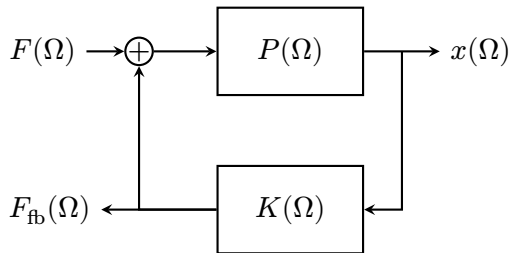
Let's return to the harmonic oscillator, which can be written very simply as

$$x(\Omega) = P(\Omega)F(\Omega),$$

where we have now collected the oscillator's force-to-displacement frequency response into the **plant** $P(\Omega)$.

Let's suppose we can devise a force $F_{\text{fb}}(\Omega) = K(\Omega)x(\Omega)$ that is a filtered version of the oscillator displacement, and apply it to the oscillator.

We call $K(\Omega)$ the **controller**.



Control alters the system dynamics

Feedback alters our oscillator's frequency-domain equation to

$$x(\Omega) = P(\Omega)[F(\Omega) + F_{\text{fb}}(\Omega)] = P(\Omega)[F(\Omega) + K(\Omega)x(\Omega)].$$

Solving for x ,

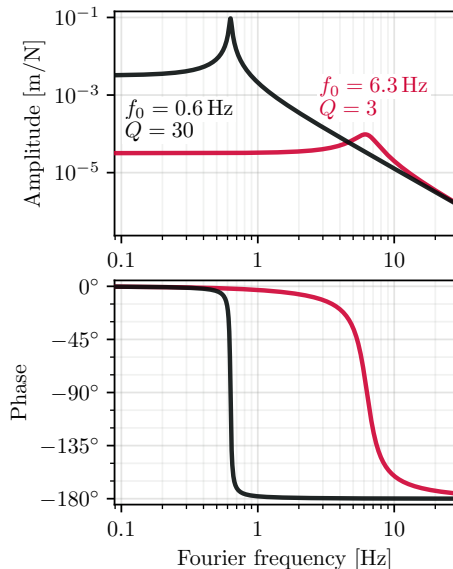
$$x(\Omega) = \frac{P(\Omega)}{1 - P(\Omega)K(\Omega)}F(\Omega).$$

We find that the system's bare plant P has been altered to $P/(1 - PK)$.

A proportional-velocity controller

Let's choose $F_{\text{fb}}(\Omega) = (-m\Omega_{\text{fb}}^2 + im\Gamma_{\text{fb}}\Omega)x(\Omega)$.
Then

$$\frac{P}{1 - PK} = \frac{1/m}{\Omega_0^2 + \Omega_{\text{fb}}^2 - \Omega^2 + i(\Gamma + \Gamma_{\text{fb}})\Omega}.$$



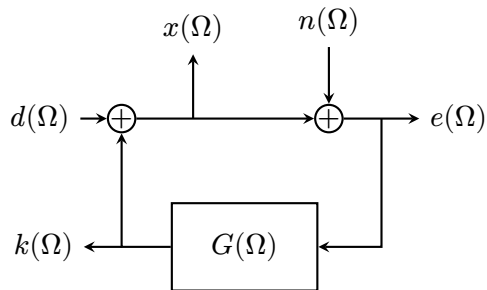
The dimensionless frequency response $G(s) = P(s)K(s)$ is the **open-loop gain**.

Two closely related quantities

$$\frac{1}{1 - G(s)} \quad \text{and} \quad \frac{G(s)}{1 - G(s)}$$

are both called **closed loop gain** in various contexts (also “sensitivity” and “complementary sensitivity”, which we will not use).

A canonical control loop with noise



The error signal is given by

$$e(\Omega) = \frac{1}{1 - G(\Omega)} [d(\Omega) + n(\Omega)]$$

When the loop gain is large,
 $e(\Omega) \simeq -[d(\Omega) + n(\Omega)]/G(\Omega)$: the external disturbance is suppressed.

The control signal is given by

$$k(\Omega) = \frac{G(\Omega)}{1 - G(\Omega)} [d(\Omega) + n(\Omega)]$$

When the loop gain is large,
 $k(\Omega) \simeq -[d(\Omega) + n(\Omega)]$: the control signal cancels the external disturbance.

The frequency response of a feedback loop to external disturbances depends fundamentally on $1/(1 - G)$.

Like any other linear system, this frequency response is only stable if the poles of $1/(1 - G)$ are all in the left-hand plane.

Alternately, one can see by inspection that a sufficient (but not necessary) criterion for instability is the **Bode criterion**: if $G(s) = +1$ for some frequency s , then a finite disturbance at that frequency produces an infinite response at that frequency.

Note, this is a complex-valued criterion: instability occurs only if $|G(s)| = 1$ and $\arg G(s) = 0$.

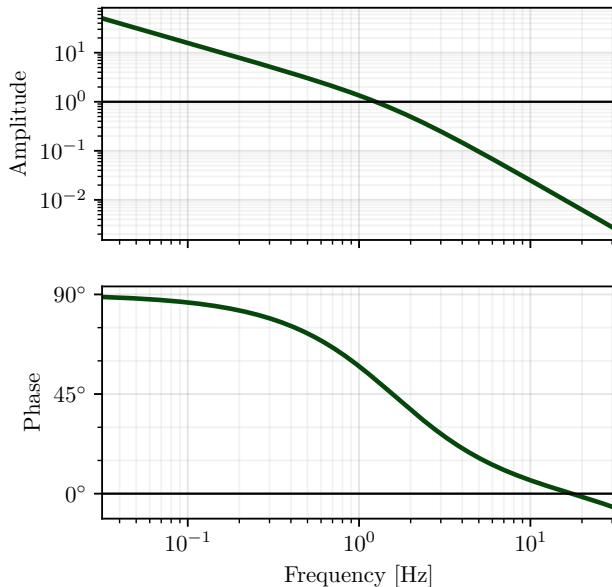
Unity gain frequency, and gain and phase margin

The frequency $f_\star$ for which $|G(f_\star)| = 1$ is the **unity-gain frequency**.

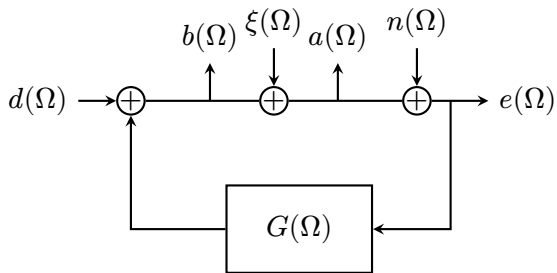
The phase $\arg G(f_\star)$ at the unity gain frequency is the loop's **phase margin**. It characterizes how much additional phase delay could be added to the loop before it becomes unstable.

For any f such that $\arg G(f) = 0$, the magnitude $|G(f)|$ constitute a **gain margin** for the loop. It characterizes how much the gain magnitude can be changed before the loop becomes unstable.

An example open-loop gain



Transfer function measurement



By injecting an excitation $\xi(\Omega)$ and measuring before $[b(\Omega)]$ and after $[a(\Omega)]$, we can estimate $G(\Omega)$:

$$\begin{aligned}\hat{G} &= \frac{\langle \xi^* b \rangle}{\langle \xi^* a \rangle} \\ &= \frac{G \langle \xi^* \xi \rangle + \langle \xi^* d \rangle + G \langle \xi^* n \rangle}{\langle \xi^* \xi \rangle + \langle \xi^* d \rangle + G \langle \xi^* n \rangle} \\ &\simeq G\end{aligned}$$

if ξ is sufficiently strong.